# MSSM Electroweak Baryogenesis and Flavour Mixing in Transport Equations

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#### Outline

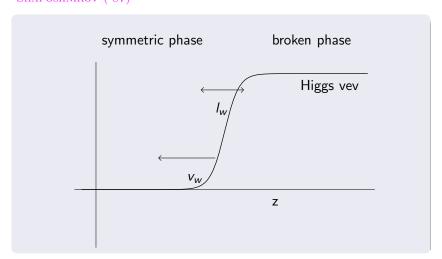
- Introduction
  - Electroweak Baryogenesis
- Pormer Approaches
  - Summary
- 3 First Principle Approach
  - The Kadanoff-Baym Equations
  - Results
- Conclusions

## Introduction

## Statement I

Electroweak Baryogenesis

= Transport + CP Violation



# CP-violation: Chargino masses in the MSSM

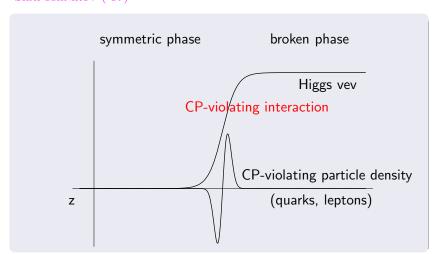
In the MSSM case the higgsino/wino mass matrix is:

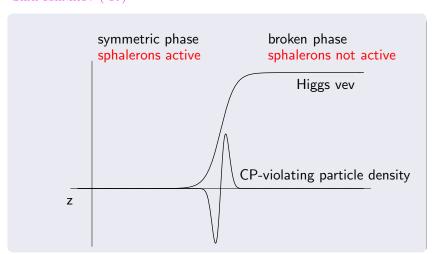
$$m = \left(\begin{array}{cc} M_2 & g h_2(z) \\ g h_1(z) & \mu_c \end{array}\right)$$

with  $M_2$  and  $\mu_c$  containing a CP-violating complex phase.

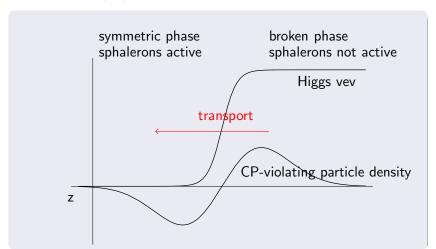
#### Qualitative difference in CP-violation

- $(\partial_z m m^{\dagger} m \partial_z m^{\dagger}) \sim \text{Im}(M_2 \mu_c^*) \neq 0$  and CP-violation is already present on the tree level.
- In the SM:  $\partial_z m \sim m \to (\partial_z m m^\dagger m \partial_z m^\dagger) = 0$  and CP-violation depends on loop calculations.





# Picture of Electroweak Baryogenesis



# Former Approaches

#### Statement II

Weaknesses of Former Approaches based on Classical Transport

A) Flavour Basis Invariance

B) Ambiguities

Particle distribution function of species j:  $f_j(\vec{p}, \vec{x}, t)$  Boltzmann-Equation:

$$\left(\partial_t + \frac{\vec{p}}{m} \cdot \nabla + \frac{\vec{F}_j}{m} \cdot \partial_{\vec{p}}\right) f_j = \text{gain- and loss-terms}$$

## Summary

# Summary of former Approaches

| Approach            | Cline et al.        | Carena et al.        |
|---------------------|---------------------|----------------------|
| <b>CP-violation</b> | dispersion relation | local source term    |
|                     | WKB                 | perturbation theory  |
| basis               | mass eigenbasis     | flavour eigenbasis   |
| quasiparticles      | charginos           | higgsinos/winos      |
| transport           | classical           | classical            |
|                     | Boltzmann type      | diffusion            |
| mixing              | not included        | in the source        |
|                     |                     | not in the diffusion |
| $\hbar$ order       | second order        | first order          |

To decide which approach is correct one has to derive semi-classical transport equations from first principles.



## Quantum Transport

#### Statement III

Quantum Transport Equation

 Kadanoff-Baym-Equation of the Wightman Function in Wigner Space Introduction

## The Wightman Function

In statistical QFT all two-point functions obtain an additional  $2 \times 2$  structure from the *in-in-*(closed-time-path)-formalism

$$\Sigma = \begin{pmatrix} \Sigma^t & \Sigma^> \\ \Sigma^< & \Sigma^{\overline{t}} \end{pmatrix}, S = \begin{pmatrix} S^t & S^> \\ S^< & S^{\overline{t}} \end{pmatrix}.$$

First Principle Approach

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In Fourier space the Greens function entry  $S^{<}$  is called Wightman function and can interpreted as particle density, e.g. in thermal equilibrium

$$iS_{eq}^{<} = 2\pi \operatorname{sign}(k_0) \, \delta(k_u^2 - m^2) \, n_{BE}(k_0)$$

such that

$$\int_{k_0>0} \frac{dk_0}{2\pi} \, 2ik_0 \, S_{eq}^{<} \sim n_{BE}(\sqrt{\vec{k}^2 + m^2})$$



Introduction

## Dynamics of the Wightman Function

The dynamics of the Wightman function is described by the Kadanoff-Baym equations that are the statistical analogue to the Schwinger-Dyson equations:

First Principle Approach

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$$\int d^4 z \left( S_0^{-1}(x_{\mu}, z_{\mu}) - \Sigma(x_{\mu}, z_{\mu}) \right) S(z_{\mu}, y_{\mu}) = \mathbb{1} \delta(x_{\mu} - y_{\mu})$$

Since the Greens-function depends not only on the relative coordinate, but also on the average coordinate  $X_{\mu} = (x_{\mu} + y_{\mu})/2$ , Fourier-transformation leads to the Moyal star product in Wigner space:

$$e^{-i\lozenge}\{S_0^{-1}(X_{\mu},k_{\mu})-\Sigma(X_{\mu},k_{\mu})\}\{S(X_{\mu},k_{\mu})\} = 1$$

with the diamond operator

$$2\Diamond\{A,B\} := \partial_{X^{\mu}}A\,\partial_{k_{\mu}}B - \partial_{k_{\mu}}A\,\partial_{X^{\mu}}B$$

First Principle Approach 000000000000

Kandanoff-Baym-Equations

# **Gradient Expansion**

## Statement IV

Semi-Classical Expansion

= Gradient Expansion in  $\Diamond$ 

Introduction

## Gradient Expansion

Since the background in the MSSM is weakly varying  $(I_w \approx 20/T_c)$  the Moyal star product can be simplified by the semi-classical approximation

$$\partial_k \partial_X pprox rac{1}{T_c I_w} pprox rac{1}{20} \ll 1 \quad o \quad e^{-i \Diamond} pprox 1 - i \Diamond$$

First Principle Approach

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Neglecting selfenergies, the simplest example for an transport equation in a varying background is for one bosonic flavour with real z-dependent mass

$$(k^{2} - m^{2}(z))S^{<} = 0$$
$$(k^{\mu}\partial_{\mu} - \frac{1}{2}(\partial_{z}m^{2}(z))\partial_{k_{z}})S^{<} = 0.$$

# Fermionic Systems

After spin projection the fermionic system of equations reads

where  $S_0 \dots S_3$  are  $2 \times 2$  matrices in flavour space and s denotes the spin.

## Transport in the Chargino sector

The chargino transport equations for the left/right handed CP violating deviations from equilibrium ( $\delta S^<=S^<-S^<_{eq}$ ) are of the form

$$k_0 \partial_t \delta S^{<} + k_z \partial_z \delta S^{<} + \frac{i}{2} [m^2, \delta S^{<}] + \text{Force}(\delta S^{<}) = \text{Source}(S_{eq}^{<})$$

#### Comments

- The term  $\frac{i}{2}[m^2, \delta S^<]$  will lead to an oscillatory behaviour of the off-diagonal particle densities, similar to neutrino oscillations with frequency  $\sim (m_1^2 m_2^2)/k_z$ .
- Without this oscillation term the first order mixing terms will not lead to CP-violation

| Approach       | Heidelberg                                  |  |
|----------------|---|--|
| technique      | Kadanoff-Baym equations                     |  |
| basis          | basis independent                           |  |
| quasiparticles | no quasi-particles / no dispersion relation |  |
| transport      | semi-classical                              |  |
|                | diffusion                                   |  |
| mixing         | included                                    |  |
| $\hbar$ order  | first order and second order                |  |
| comment        | accounts for oscillations                   |  |
|                | unambiguous                                 |  |

## Oscillation Effects

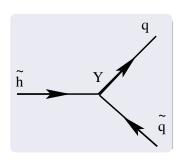
## Statement V

The Flavour Oscillation Leads to A) Suppression of Mixing far from Degeneracy B) BAU of maximally  $4 \times \eta_{exp}$ 

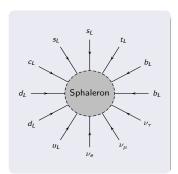
#### Determination of the BAU

#### HUET, NELSON ('95)

The missing parts to determine the baryon asymmetry of the universe are:

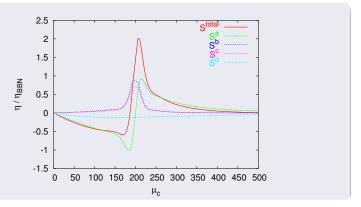


and



Results

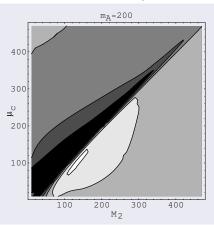
Parameters chosen:  $M_2 = 200$  GeV,  $v_w = 0.05$ ,  $I_w = 20/T_c$ , CP-phase maximal.



## Numerical Results II

Results

Parameters chosen:  $v_w = 0.05$ ,  $I_w = 20/T_c$ , CP-phase maximal.



## Pessimistic Conclusion

MSSM electroweak baryogenesis is a rather unlikely scenario based on

- A light stop to acquire a strong first order phase transition
- The condition  $\mu_c \approx M_2 \lesssim 400$  GeV of the *a priori* unrelated parameters  $M_2$  and  $\mu_c$
- A large CP-violating phase that hardly satisfies experimental EDM bounds

## Optimistic Conclusion

We achieved to attain a formalism that

- Describes semi-classical transport
- Is basis independent
- Is derived from first principles
- Unambiguous

and it unveiled an oscillation in the dynamics that is significant for the first order mixing effects.

Based on our analysis, MSSM electroweak baryogenesis is a scenario that

- Can explain the BAU
- Predicts  $\mu_c \approx M_2 \lesssim 400 \text{ GeV}$
- Will be testable soon



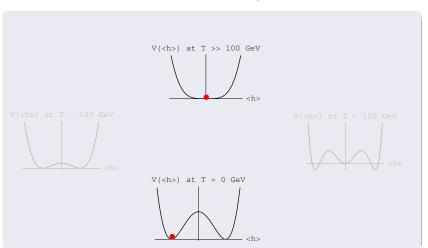
#### Conclusion

#### Statements I-V

- I) Electroweak Baryogenesis = Transport + CP Violation
- II) Weaknesses of Former Approaches based on Classical Transport
  - Basis Invariance
  - Ambiguities
- III) Quantum Transport Equation = Kadanoff-Baym-Equation of the Wightman Function in Wigner Space
- IV) Semi-Classical Expansion = Gradient Expansion
- V) The Flavour Oscillation Leads to
  - Suppression of Mixing far from Degeneracy
  - BAU of maximally  $4 \times \eta_{exp}$

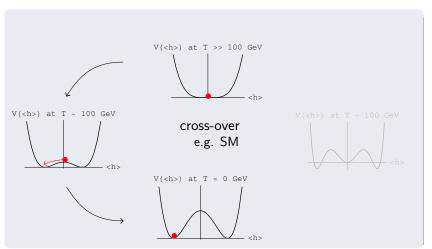
## Non-equilibrium: The electroweak phase transition

Cross-over versus first order electroweak phase transition:



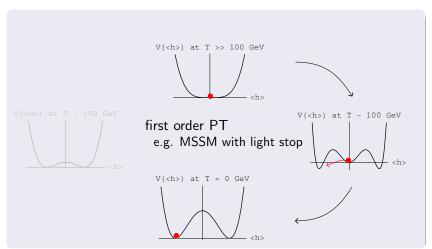
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## Non-equilibrium: The electroweak phase transition

Cross-over versus first order electroweak phase transition:



#### Sakharov conditions

Baryogenesis is one of the cornerstones of the Cosmological Standard Model.

The celebrated Sakharov conditions state the necessary ingredients for baryogenesis:

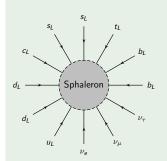
#### Sakharov conditions

- C and CP violation
- B-violation
- non-equilibrium

## C- and B-violation: The sphaleron process

In the hot universe B- and C-violation is present in the SM and its extensions due to sphaleron precesses.

#### The effective sphaleron vertex



- $\Delta B = 3$ ,  $\Delta L = 3$ ,  $\Delta N_{CS} = 1$
- B L conserving
- B + L violating
- Exponentially suppressed by the W mass
- Topological effect of the SU(2) gauge sector

## Former Approach using Dispersion Relations

CLINE, JOYCE, KAINULAINEN ('97, '00)

The CP violating dispersion relation E(p,z) for a varying mass (e.g.  $m(z) = |m(z)|e^{i\theta(z)}$ ) is determined by the WKB method

$$E^2 = \vec{p}^2 + m^2 \pm \frac{m^2 \partial_z \theta}{2p_z}$$

and put into the classical Boltzmann equation for the particle density f:

$$(\partial_t + \partial_{p_z} E \partial_z - \partial_z E \partial_{p_z}) f = \text{Coll.}$$

This procedure is performed in mass eigenbasis and does not contain mixing effects.

## Former Approach using a Local Source

CARENA, MORENO, QUIROS, SECO, WAGNER ('00, '02)

The additional term in the Schwinger-Dyson equation is interpreted as an interaction term

The CP violating source is calculated from

$$j^{\mu}(X_{\mu}) = \frac{1}{\tau} \int \frac{d^4p}{(2\pi)^4} p^{\mu} \delta S^{<}(p_{\mu}, X_{\mu})$$

and put 'by hand' into diffusion equations

$$(D\partial_z^2 - v_w \partial_z - \Gamma)f = j^0.$$

The source  $j^{\mu}$  is determined in flavour eigenbasis and contains mixing.



## Status of Electroweak Baryogenesis

#### EWB in the SM excluded

- CP violation too small
- No strong first order phase transition

#### EWB in the MSSM not excluded/disproved

- CP violation in the neutralino/chargino sector
- Strong first order phase transition in case of a light stop