

Photon Mixing in Domain Walls and Cosmic Coincidences

**Malcolm Fairbairn
Stockholm University**

hep-ph/0507020

The Coincidence Problem

- Observations of Type 1a supernovae suggest $\Omega_{\Lambda}=0.7$, $\Omega_M=0.3$
- That means $\rho_{\Lambda}=\rho_M$ at $z=0.3$
- Does this need to be explained?
- Quintessence models exist, but require fields with masses $m_Q=10^{-33}\text{eV}=(\text{Hubble radius})^{-1}$



Anthropic Arguments:

If ρ_{Λ} was even slightly larger, galaxies could not form (Martel, Shapiro & Weinberg astro-ph/9701099) and life could not exist



**So is it time to give up
the scientific method?**

**Or can we think of
an alternative?**



Photon-Axion Mixing

Csaki, Kaloper & Terning hep-ph/0111311

ultra light axion, photons from supernovae convert into axions in the intergalactic magnetic field, gives the illusion of $\Lambda > 0$

$$L = -\frac{1}{2}(\partial^\mu a \partial_\mu a + m^2 a^2) + \frac{a}{M} F_{\mu\nu} \tilde{F}^{\mu\nu} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

probability of a photon remaining a photon:-

$$P_{\gamma \rightarrow \gamma} = 1 - \frac{B^2 \omega^2}{m^4 M^2 + B^2 \omega^2} \sin^2 \left[\frac{\sqrt{m^4 M^2 + B^2 \omega^2}}{\omega M} L \right]$$

to get right dimming, we require $M/B = \text{Hubble Length}$

to suppress oscillations of CMB photons:-

$$\omega_{\text{optical}} > \frac{m^2 M}{B} > \omega_{\text{cmb}}$$

intergalactic electrons cause problems

Mixing with Paraphotons

$$L = -\frac{1}{4g^2} F^{\mu\nu} F_{\mu\nu} - \frac{1}{4g'^2} G^{\mu\nu} G_{\mu\nu} - \frac{\chi}{g'g} F^{\mu\nu} G_{\mu\nu}$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

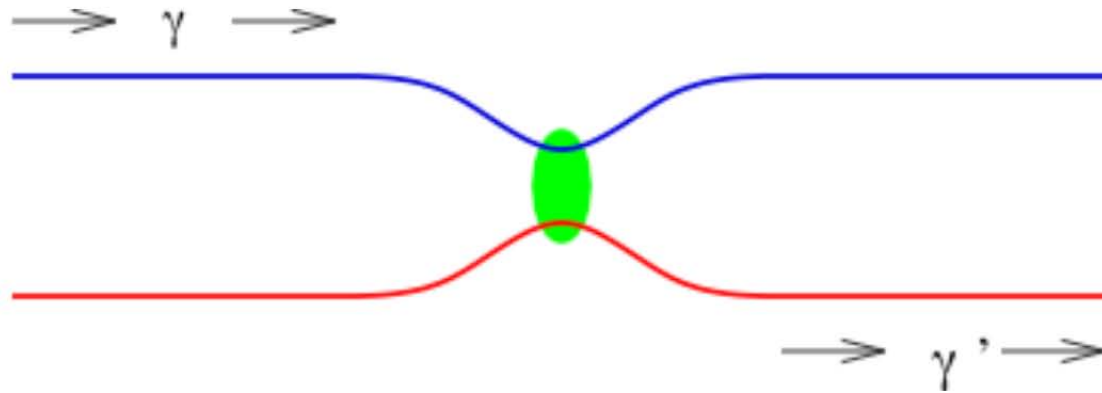
$$G_{\mu\nu} = \partial_\mu A'_\nu - \partial_\nu A'_\mu$$

Additional U(1) gauge field weakly coupled to electromagnetism ($\chi \ll 1$) except inside the core of domain walls

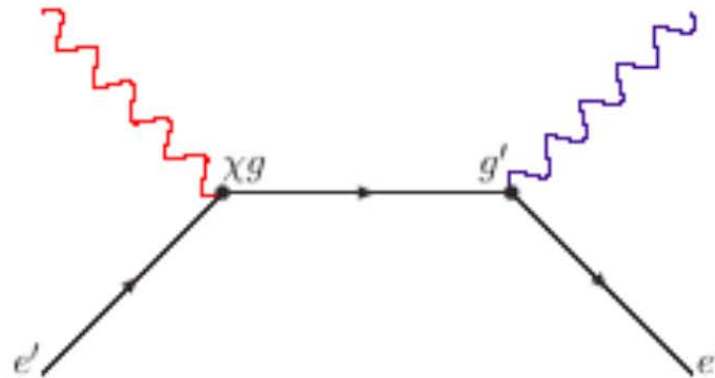
hep-ph/0507020

Photon Mixing in Domain Walls

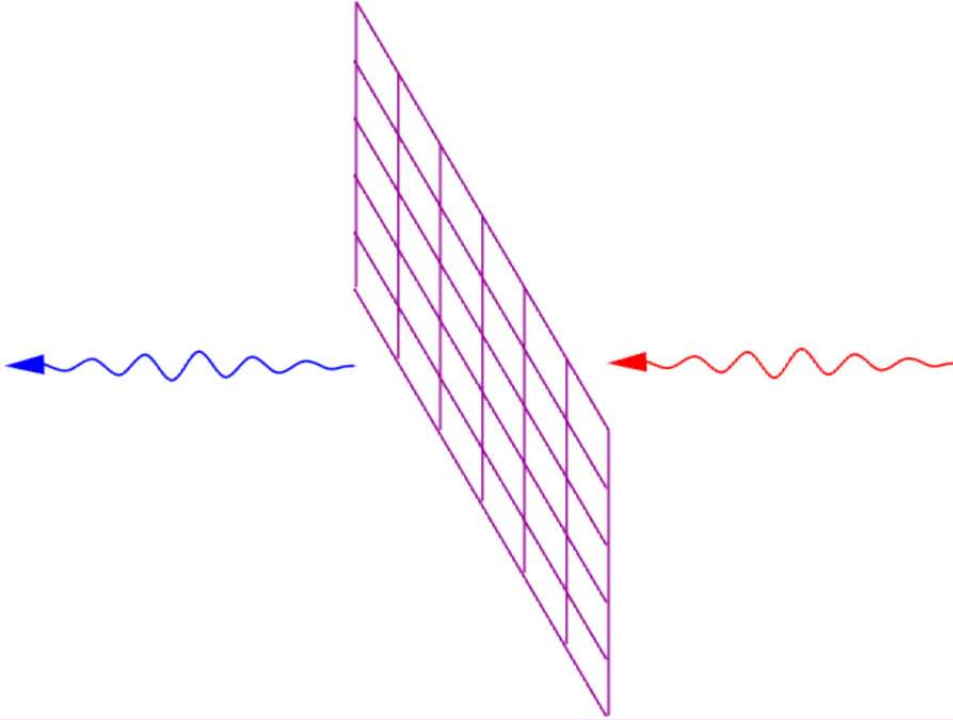
kinks in relative position of branes in compact space



Inside core, branes closer than mass of fields in bulk



Photons and paraphotons massless :-
need population of paraelectrons to introduce refractive index.

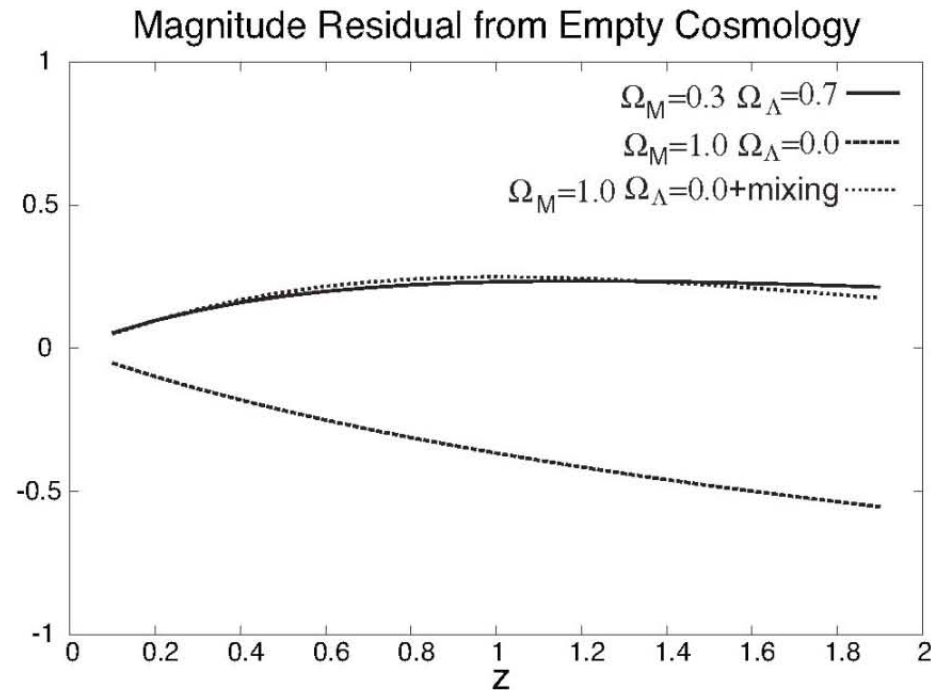


Since there is one domain wall per horizon, approximately 50% of photons created at redshift $z=1$ are lost. This does not depend on any of the energy scales in the model.

$$P_{\gamma \rightarrow \gamma} = \frac{1}{2} \left(1 + e^{-N(z)} \right)$$

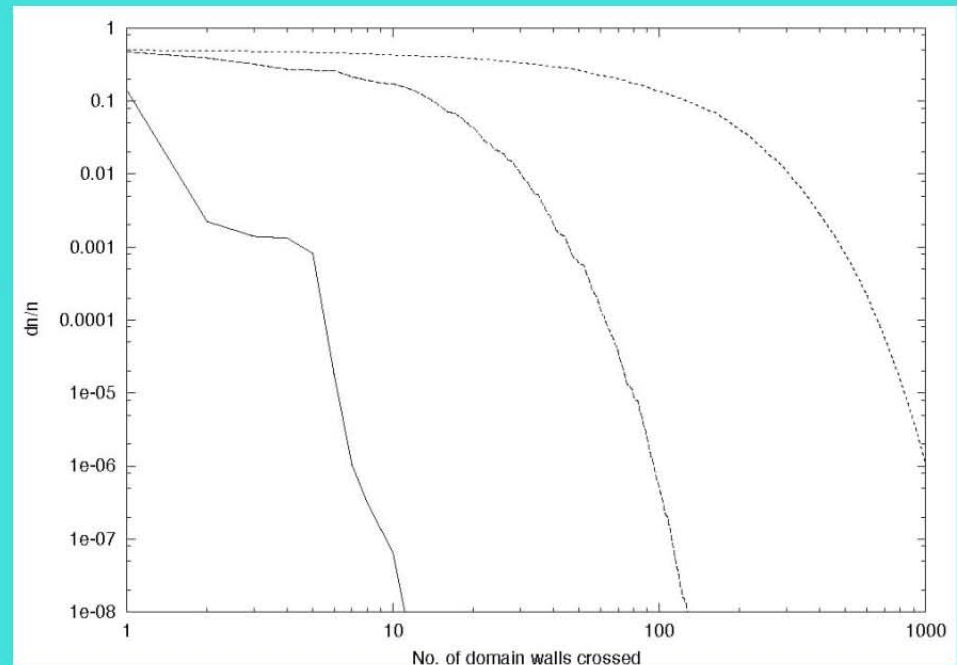
Where $N(z)$ is the number of domain walls crossed by a photon emitted at redshift z

hep-ph/0507020



Isn't this dangerous for CMB photons ?

- CMB photons reaching earth today have crossed ~ 1000 domain walls
- Distortion due to mixing of photons less than 10^{-5} rather quickly ~ 10 walls for 50% probability of mixing per wall
- Para-electrons need to be cool to avoid SZ effect while photons are on other brane



hep-ph/0507020

Conclusions

- Photon mixing in Domain walls changes luminosity distance
- No mass scale set to inverse Hubble length required
- Observers at any redshift conclude universe has recently started to accelerate

hep-ph/0507020