Moduli Stabilization in Heterotic Flux Compactifications

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Outline

- Introduction: Compactifications, moduli and flux
- Manifolds with SU(3) structure
- Heterotic string and SU(3) structure
- Half-flat mirror manifolds
- Generalized half-flat manifolds
- Conclusions and outlook

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Based on: hep-th/0507173 and hep-th/0408121

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In collaboration with: B. de Carlos (Sussex), S. Gurrieri (Ky-

oto), A. Micu (Sussex)

Compactification of D = 10 theory

Compactify on d=6 space with topology X

Metric: $ds^2=dx^\mu dx^\nu \eta_{\mu\nu}+g_{mn}(y,b^i)dy^m dy^n$, where $g_{mn}(b^i)$ are metrics on X with moduli b^i .

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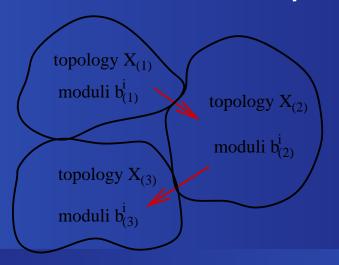
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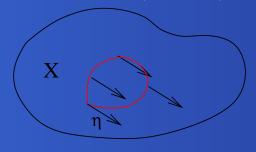
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Schematic structure of moduli space:



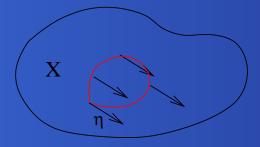
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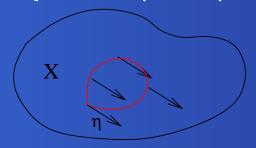
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Introduce forms:

$$J_{mn} = \eta^{\dagger} \gamma_{mn} \eta$$

 $\Omega_{mnp} = \eta^{\dagger} \gamma_{mnp} \eta$

Kahler form complex structure

Calabi-Yau spaces, cont'd

On a Calabi-Yau space

$$\nabla J = 0 \Rightarrow dJ = 0$$

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Ricci-flat metric in one-to-one correspondence with

 $J=t^i\omega_i$ and $\Omega=Z^A\alpha_A-\mathcal{G}_A\beta^A$, where $\omega_i,\,i=1,\ldots,h^{1,1}$ are two-forms and $\alpha_A,\,\beta^A,\,A=0,\ldots,h^{2,1}$ are three-forms. Note: $d\omega_i=d\alpha_A=d\beta^A=0$.

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Kahler moduli complex structure moduli dilaton

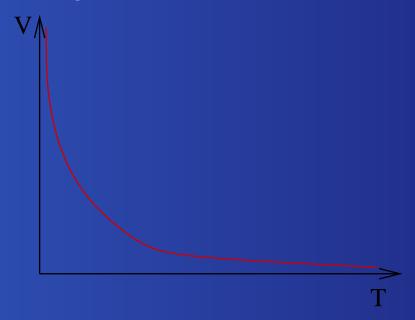
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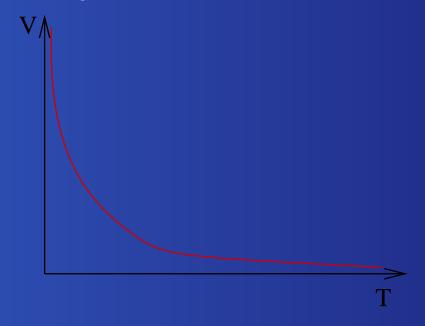
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Hard to fix all moduli and potential very steep.

Flux

String theory contains p-forms $A_{I_1...I_p}$ with field strengths F = dA.

Flux: Internal components of F are non-zero, $F_{m_1...m_{p+1}} \neq 0$.

Quantization: $\int_{C \subset X} F \in \mathbf{Z}$

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 \rightarrow flux generates a D=4 moduli potential

Example IIB

Flux: NSNS 3-form H and RR 3-form F

Flux superpotential: $W_{\text{flux}} = \int_X (F - iSH) \wedge \Omega$

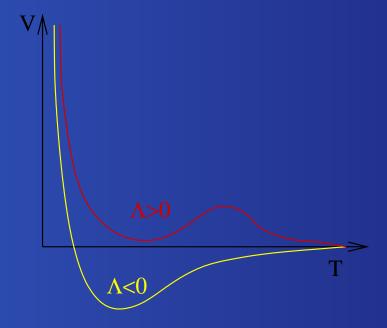
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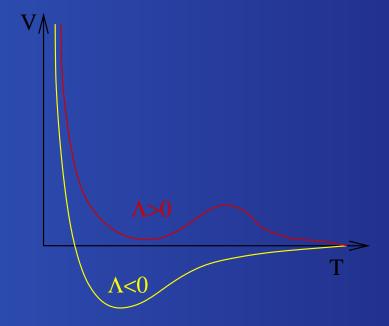


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Can fix all moduli and less steep.

Heterotic string

Why heterotic?

- Gauge unification natural
- Large number of supersymmetric models
- Allows fermions from 16 of SO(10)

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Such manifolds can be classified by the SU(3) content of the torsion κ .

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Consistency of vacuum: $|\mathcal{W}| \ll 1$ at minimum.

Half-flat mirror manifolds

Arise in the context of type II mirror symmetry.

Characterized by: (Gurrieri, Louis, Micu, Waldram, '02)

$$d\omega_i = e_i \beta^0$$
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Magnetic torsion?

Simple STZ model

Three moduli: S, T, Z

Kahler- and superpotential:

$$K = -\ln(S + \bar{S}) - 3\ln(T + \bar{T}) - 3\ln(Z + \bar{Z})$$

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Validity: Need large radius, Re(T) > 1, large complex structure, Re(Z) > 1 and weak coupling Re(S) > 1.

Reminder, D=4 supergravity

Supergravity potential:

$$V = e^K \left(K^{X\bar{Y}} F_X \bar{F}_{\bar{Y}} - 3|W|^2 \right)$$

where
$$F_X = \partial_X W + \partial_X K W$$
, $K_{X\bar{Y}} = \partial_X \partial_{\bar{Y}} K$.

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Supersymmetric vacua: $F_X = 0$

Cosmological constant: $V|_{F_X=0} = -3e^K|W|^2 \le 0$

Simple STZ model, results

- All fields stabilized, except for one combination of Im(T) and Im(Z) which is flat.
- $x \equiv c \text{Re}(S) = 1/4 \rightarrow \text{barely weak coupling}$
- difficult to achieve Re(T) > 1 and Re(Z) > 1

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- difficult to achieve Re(T) > 1 and Re(Z) > 1
- → hard to find consistent vacua

General case

Arbitrary number of moduli: S, T^i , Z^a

Superpotential: $\mathcal{W}=e_iT^i+\epsilon_aZ^a+rac{i}{2} ilde{d}_{abc}\mu^aZ^bZ^c$

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- x = 1/4 remains true
- still difficult to achieve ${\rm Re}(T)>1$ and ${\rm Re}(Z)>1$
- → difficult to find consistent, supersymmetric vacua for half-flat mirror manifolds.

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For supersymmetric vacua:

- all moduli fixed
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- still difficult to achieve ${\rm Re}(T)>1$ and ${\rm Re}(Z)>1$

Main problem: No heterotic derivation of such a model known.

Generalized half-flat manifolds

Suggested by gauging of N=2 SUGRA Characterized by: (d'Auria et al., '04)

$$d\omega_{i} = p_{Ai}\beta^{A} - q_{i}^{A}\alpha_{A}$$

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$$\mathcal{W} = (q_i^A \mathcal{G}_A - p_{Ai} Z^A) T^i + \epsilon_A Z^A - \mu^A \mathcal{G}_A$$

If dH = 0 (standard embedding) $\mu^A p_{Ai} - \epsilon_A q_i^A = 0$.

Solution method

Superpotential: $W = \mathcal{W}(T^i, Z^a) + ke^{-cS}$

Find global solution T_0^i , Z_0^a , so that $\partial_i W(T_0, Z_0) = \partial_a W(T_0, Z_0) = 0$.

Define $\mathcal{W}_0 = \mathcal{W}(T_0, Z_0)$ and require $|\mathcal{W}_0| \ll 1$.

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Define $W_0 = \mathcal{W}(T_0, Z_0)$ and require $|\mathcal{W}_0| \ll 1$.

Then a local solution exists nearby and

$$(2x+1)e^{-x} = \left| \frac{\mathcal{W}_0}{k} \right|$$

where x = c Re(S).

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Standard embedding: $\xi r - \epsilon q + \mu p - \rho e = 0$

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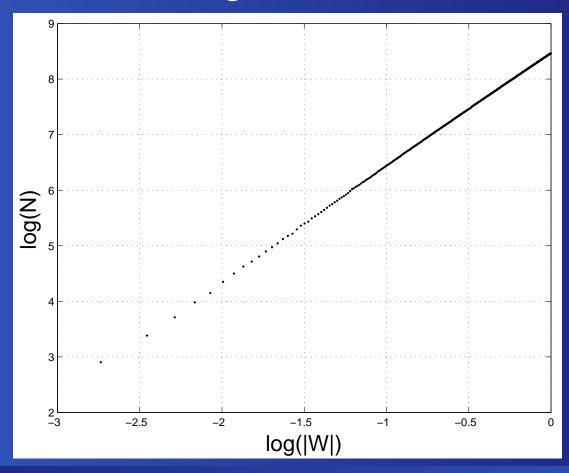
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For standard embedding: $|\mathcal{W}_0| > 0.8$ always.

Non-standard embedding

Number of global solutions as a function of $|\mathcal{W}_0|$ for parameters in range -70...70.



Non-standard embedding, cont'd

Let N(M, w) be the number of solutions for parameters in the range $-M \dots M$ and $|\mathcal{W}_0| < w$.

We find : $N(M,w) \simeq M^5 w^2$

 \rightarrow roughly 10^{-3} of vacua lead to a weak gauge coupling ${\rm Re}(S) \simeq 24$.

Conclusion and outlook

- The "missing" RR flux in the heterotic string can be "replaced" by torsion of the internal space.
- It is difficult to obtain consistent vacua for half-flat mirror spaces and for more general spaces with standard embedding.
- Generalized half-flat spaces with non-standard embedding allow for a large number of consistent vacua.
- Gauge couplings weak enough for unification are obtained for a small fraction (e. g. 1/1000) of these vacua.

Conclusion and outlook, cont'd

- The cosmological constant can probably be lifted to positive values by adding anti 5-branes.
- A number of theoretical issues need better understanding: construction of spaces, gauge field sector, quantization,...
- Analyze potentials away from vacua: cosmology, inflation?